



# A Plan for Low Energy Science at the 88-Inch Cyclotron in the Next Decade

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“Plans are useless, but planning is indispensable”.

- Dwight D. Eisenhower

“A good plan violently executed now is better than a perfect plan next week”.

- George S. Patton

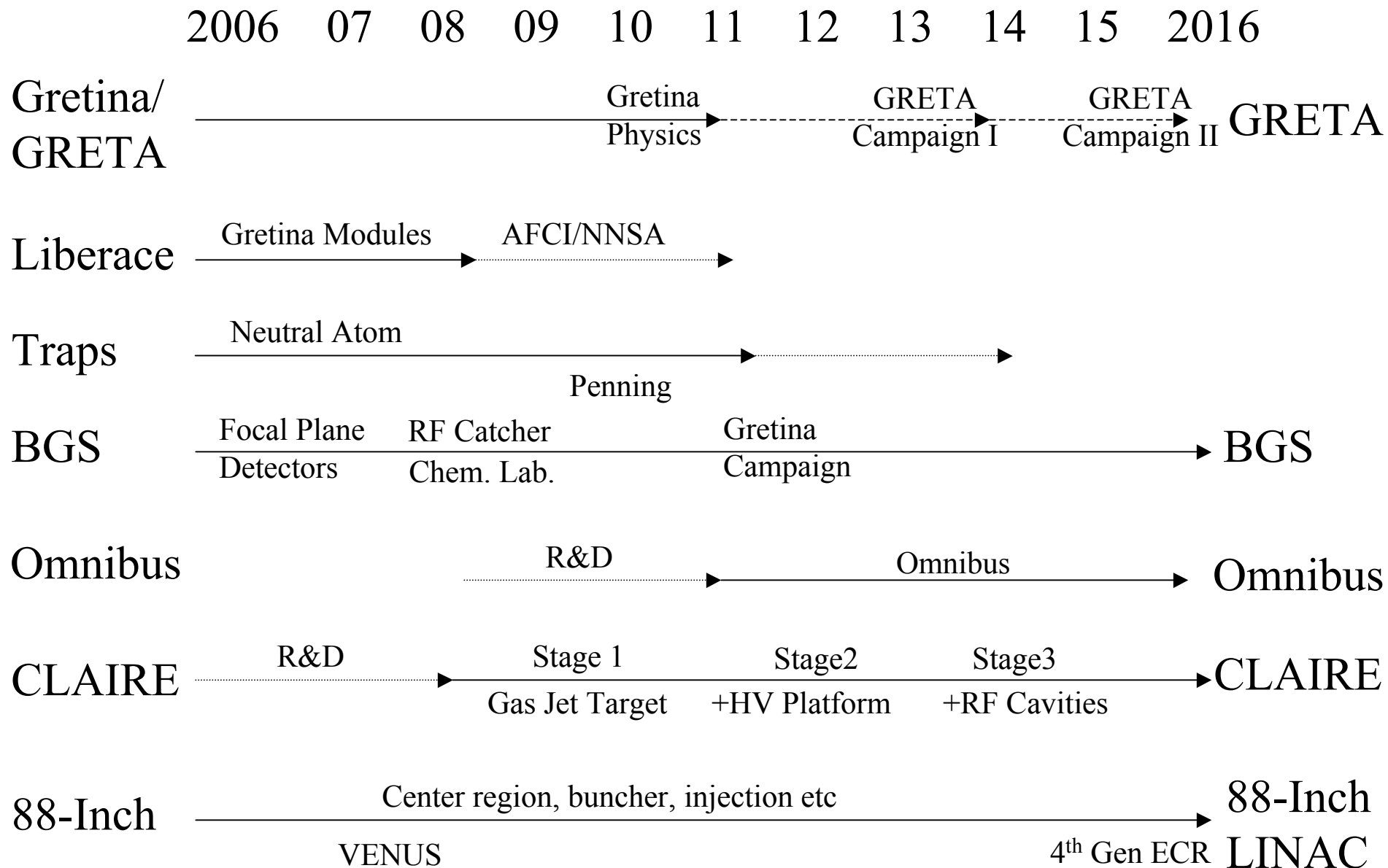




“Life is what happens to you while you are  
busy making other plans”.

- John Lennon

## A Plan for the Next Decade: Timeline



## Advantages From Operation of 88-Inch

- BASE Program
  - Vital to defense and aerospace communities
- In-house Science Programs
  - Unique heavy element research effort
  - Applied (surrogate and n-beam) work (AFCI and Defense)
  - LIBERACE science: structure, reactions, astrophysics
  - Home to fundamental symmetries program
  - Home to GRETINA/GRETA development
  - Home to VENUS ion-source development
- Training post-docs and students.
  - Connections to UCB Physics, Chemistry, and Nucl. Eng.
- Retention of broad stable beam capability.
  - Complementary role in the radioactive beam era

## In-House Low Energy Science Program: 2016

- Upgraded 88-Inch able to host scientific and applied programs including a stable-beam program with GRETA (full  $4\pi$  array).
- 88-Inch is home to two world-leading large acceptance separators, BGS and Omnibus. Unique scientific programs:
  - a) Heavy element science
    - Chemistry and atomic properties
    - Transfermium spectroscopy (prompt and delayed)
    - Mass measurements
  - b) Far-from-stability studies complementary to RIB efforts
    - Gamma-ray studies of n-rich nuclei via DIC
    - Decay studies (and tagging) of rare isotopes
- CLAIRE: National accelerator facility for nuclear astrophysics with low-energy high-intensity beams. Allied with DUSEL.

## FY07/08

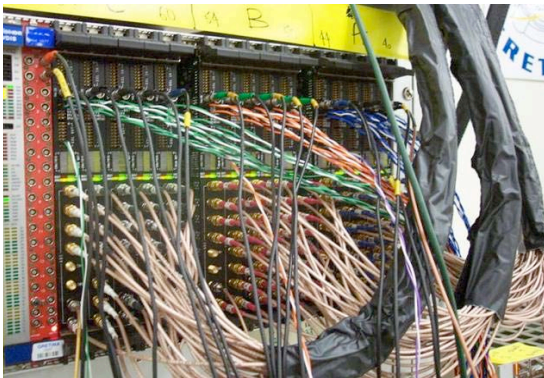
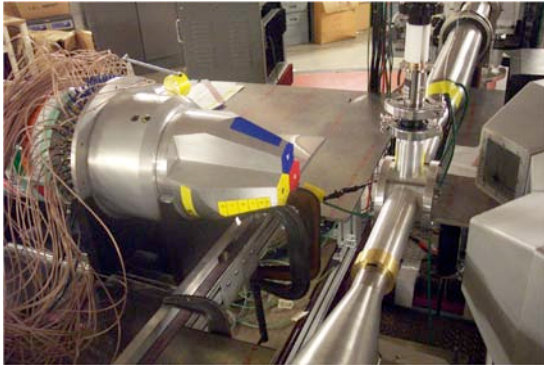
- a) *Extensive program of Nuclear Structure experiments, defense and AFCI work.*  
LIBERACE will be coupled with GRETINA modules. (LDRD funded \$300K).  
Extend Si detection to  $\sim 4\pi$  by adding Si box to existing annular set-up. \$50K.
- b) *New effort on transfermium decay spectroscopy (especially high-K isomers).*  
Focal plane detectors (especially a DSSD) for BGS. \$50K.
- c) *Extended effort on physics beyond the Standard Model.*  
New neutral atom trapping effort on  $^{18,19}\text{Ne}$ . \$150K.
- d) *New generation of heavy element chemistry experiments.*  
Pu targets and dedicated chemistry laboratory. BGS as pre-separator. \$250K.
- e) *New astrophysics program looking towards DUSEL.*  
R&D for CLAIRE. (LDRD funded \$220K).

Some examples from the above...



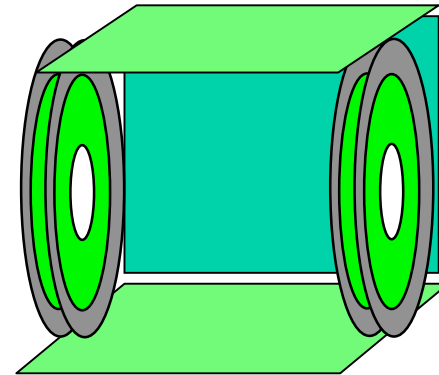
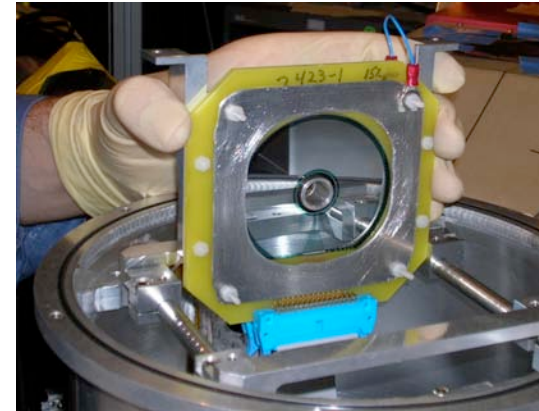
# LIBERACE Upgrades

## a) Add GRETINA Modules



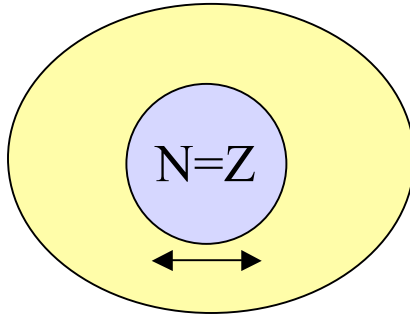
- Digital Signal Processing (DSP) for Clovers and Si detectors
- Add GRETINA Modules
- Higher  $E_\gamma$  efficiency (high- $E_\gamma$ )

## b) STARS and STRIPES

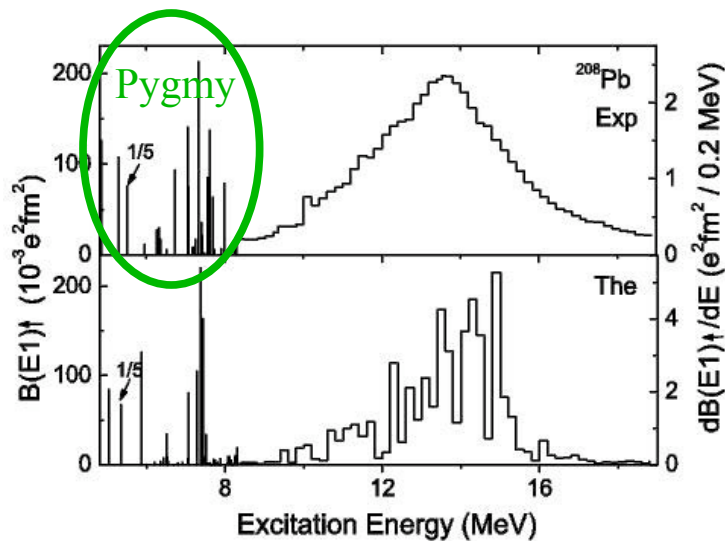


- Currently annular Si detectors (STARS)
- Add planar Si detectors box (STRIPES)
- Increased particle detection efficiency

# Fine Structure of Pygmy Resonances

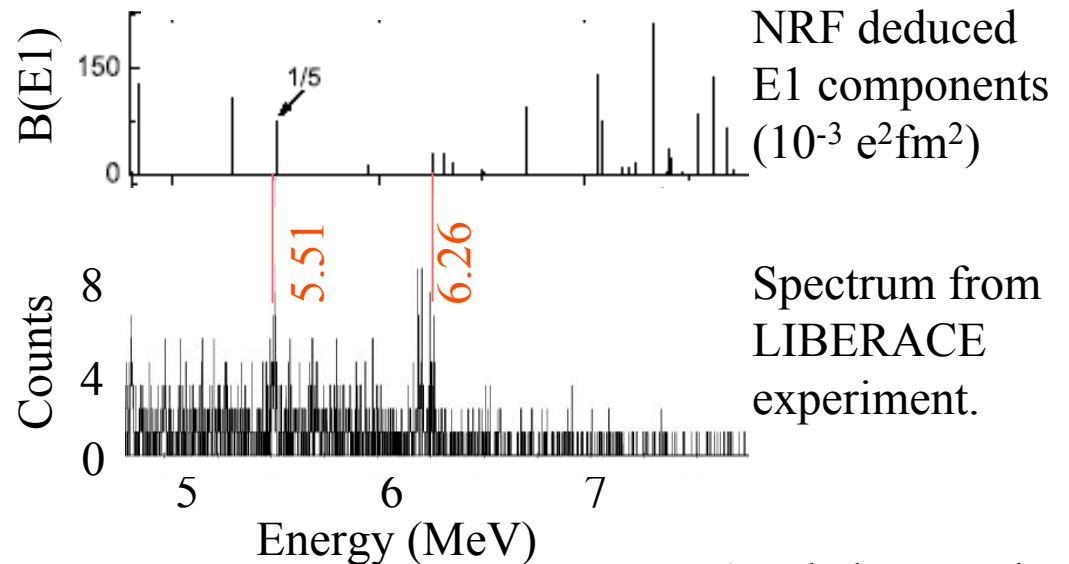


- Soft dipole mode
- Small fraction of E1 strength
- Usually studied in  $(\gamma, \gamma')$



N. Ryezayeva et al, PRL 87 272592 (2002)

- Using the  $^{208}Pb(^{17}O, ^{17}O')$  reaction with LIBERACE+STARS
- Population and decay of different states compared with  $(\gamma, \gamma')$  experiments
- Test experiments (2005/06) revealed potential of method (see below).
- Need large increase in statistics  
→ Cooling of Si detectors (> Dec 2006).  
→ **Add GRETINA modules**



M. A. Deleplanque et al.

# $^{16}\text{C}$ - Cracking the Egg

## Results from RIKEN

VOLUME 92, NUMBER 6

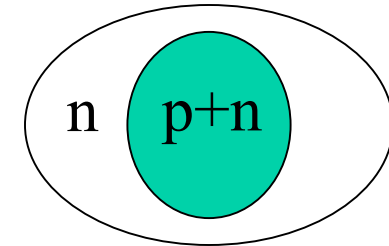
PHYSICAL REVIEW LETTERS

week ending  
13 FEBRUARY 2004

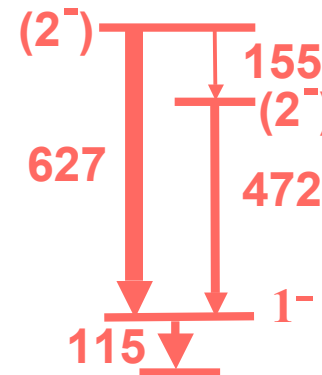
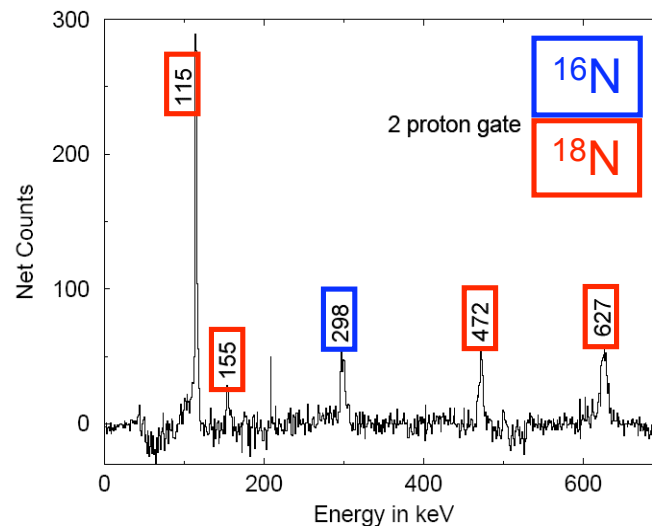
Anomalous Hindered  $E2$  Strength  $B(E2; 2_1^+ \rightarrow 0^+)$  in  $^{16}\text{C}$

PHYSICAL REVIEW C 73, 024610 (2006)

Neutron-dominant quadrupole collective motion in  $^{16}\text{C}$



- We can study  $^{16}\text{C}$  via the  $^{11}\text{B}(^7\text{Li}, 2\text{p})$  reaction with LIBERACE+STARS  
→ Measure lifetime via Doppler shift techniques and deduce  $B(E2)$
- Shown utility of 2p-evaporation studies of sd-shell light nuclei already
- $^9\text{Be}(^{11}\text{B}, 2\text{p})^{18}\text{N}$  to resolve anomalous 1- isomeric decay



M. Wiedeking et al.

$\sim 4\pi$  p-detection efficiency of STARS+STRIPES ideal for such studies

# $\beta$ - $\nu$ Correlation Measurements

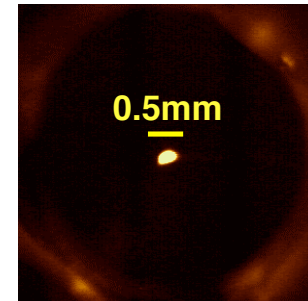
Currently: laser trapping of  $^{21}\text{Na}$  for precision electroweak tests

Laser trapped samples are:

- Cold
- Isotopically pure (contaminant-free)
- Held in vacuo: no scattering (not very dense)
- Polarizable

Ideal source for:

Full kinematic detection of a beta decay



$\sim 800,000$   $^{21}\text{Na}$  in a MOT

Beyond Standard Model Physics search:

- Non “Vector minus Axial Vector” couplings
- New Scalar Bosons
- Second class currents from quark mass hierarchy
- Massive sterile neutrinos
- Left-Right symmetric models
- Sleptons...



Lasers and nuclear physics...

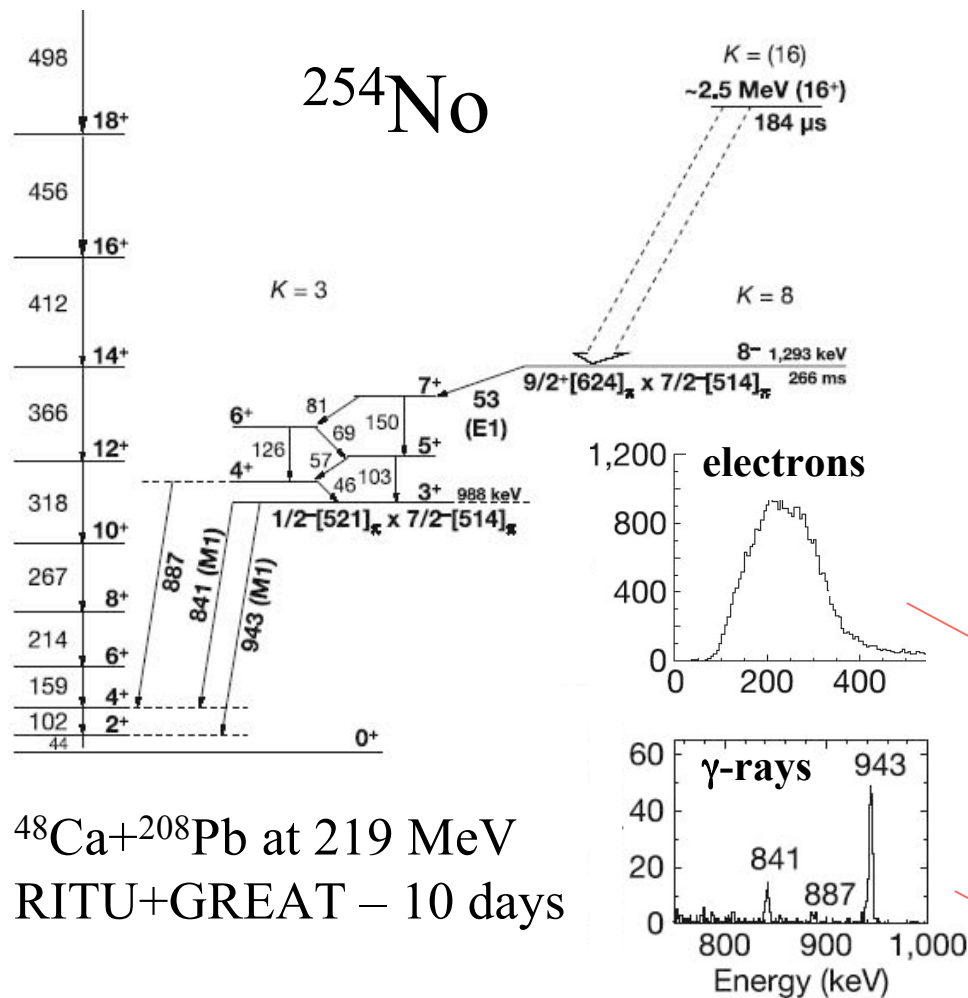
Future:  $^{18,19}\text{Ne}$  “cleanest” test for scalar or tensor currents outside Standard Model

# BGS Focal Plane Studies: K-Isomers in Transfermium Nuclei



## Nuclear isomers in superheavy elements as stepping stones towards the island of stability

R.-D. Herzberg et al., Nature 442 (2006) 896, S. Tandel et al., Phys. Rev. Lett. 97 (2006) 082502



$^{48}\text{Ca} + ^{208}\text{Pb}$  at 219 MeV  
RITU+GREAT – 10 days

- BGS superior heavy element separator to either FMA or RITU
- Clean tag on isomer decays from conversion electrons in a DSSD.
- Clovers/LEPS detect gammas.

### BGS Rates

Target = 500  $\mu\text{g}/\text{cm}^2$

Current = 0.7  $\mu\text{A}$

Chopping factor = 0.5

Efficiency (BGS)=0.6

Cross section = 2  $\mu\text{b}$

→ 6 x 10<sup>5</sup> recoils at  
focal plane /day

Fraction ms-isomer=0.5

Efficiency (DSSD)>0.5

Efficiency (Ge)>0.05

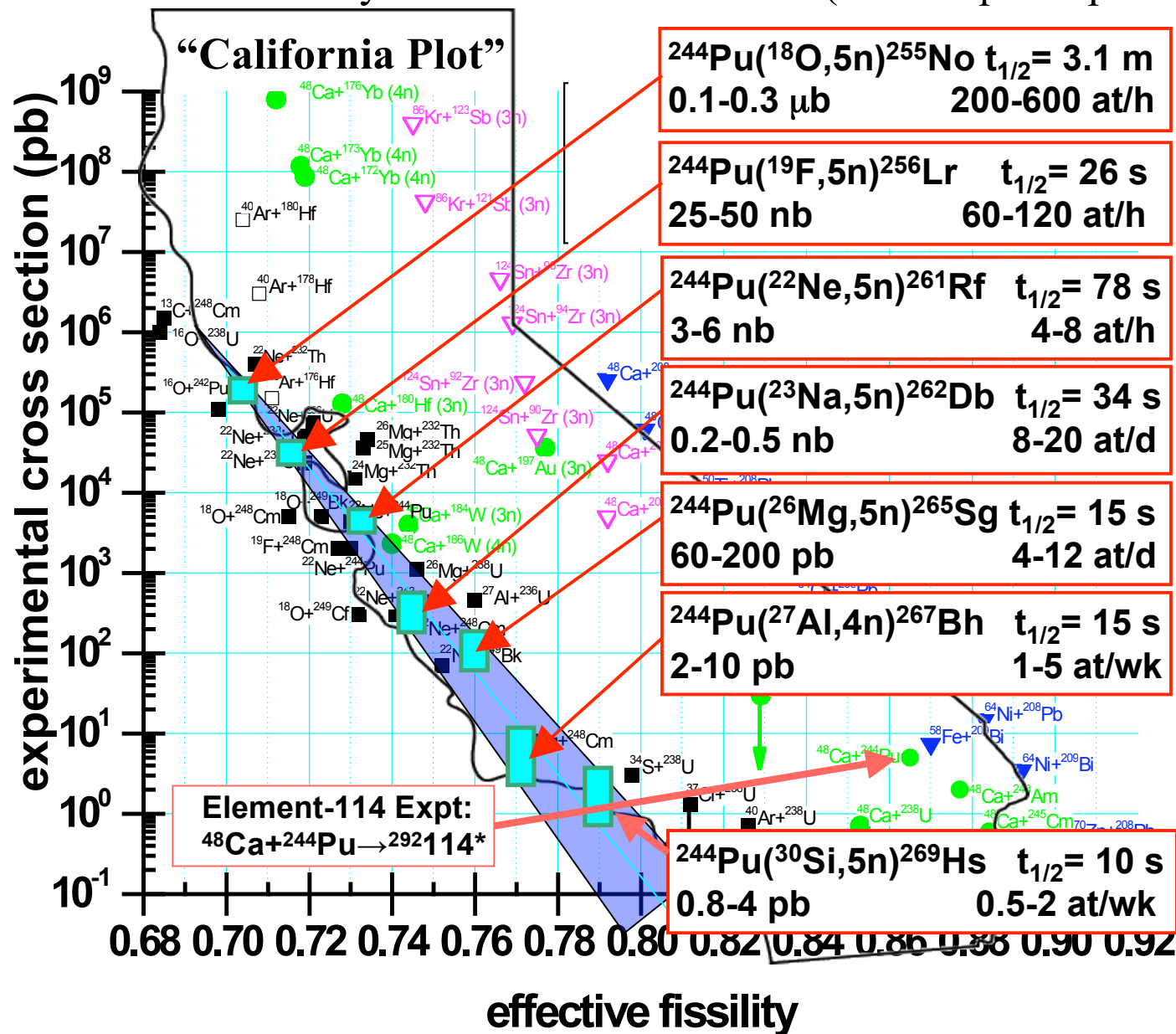
→ >7500  $\gamma$ 's/day



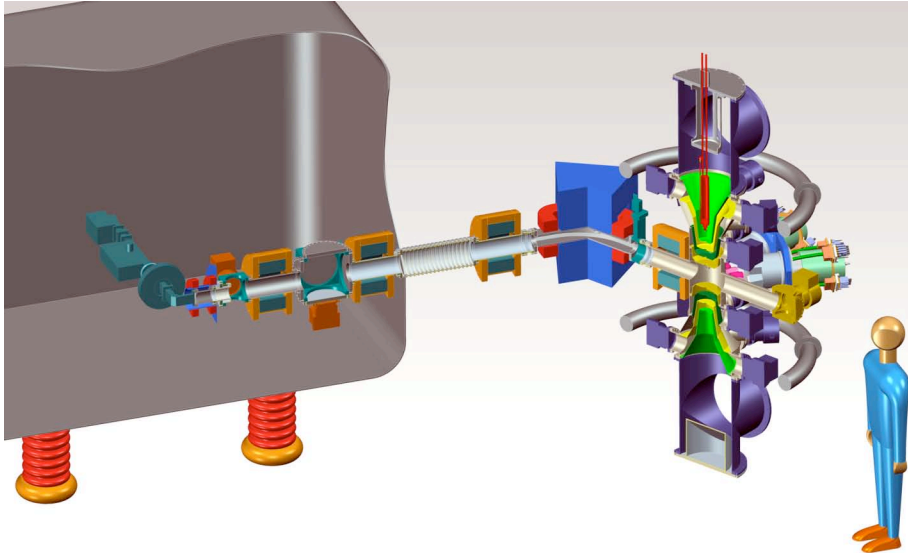
# Heavy Element Experiments with Pu Targets

Pu targets allow access to longer-lived neutron-rich isotopes

A dedicated chemistry lab is needed near BGS (used as pre-separator)



# CLAIRE: A National Accelerator for Nuclear Astrophysics



## **The Stages of CLAIRE**

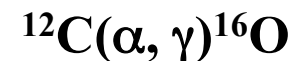
- 1) *Gas jet target* ( $\sim 10^{18-19}$  particles/cm<sup>2</sup>).  
Beams from 88-Inch to measure “ $\gamma$ -process” reactions: e.g.,  $^{96}\text{Ru}(\alpha, \gamma)$ ,  $^{124}\text{Xe}(\alpha, \gamma)$ ,  $^{136}\text{Ce}(\alpha, \gamma)$
- 2) *Add HV Platform*  
 $1^+$  ion source, analyzing magnet, focusing optics.  
Provide intense light-ion beams for measurement of key H and He burning reactions.
- 3) *Add RF Cavities and ECR Source*  
Reactions in CNO and Ne-Na-Al Cycles.

- Meet the need identified in last LRP of a “...dedicated, high intensity ( $\sim$ mA), low-energy ( $<3$  MV), accelerator..”
- Required in order to understand stellar evolution especially H and He burning.

## **Two examples of key reactions**



- Crucial to understanding of solar neutrino flux.
- Needed to 1% accuracy in region of Gamow window



- Determines fate of intermediate mass stars (black hole vs. neutron star).
- Fixes C/O ratio which influences all later nuclear burning stages.

Strong Connection with DUSEL (proposal for gas-jet target)

## FY09/10

a) *Decay spectroscopy in low background environment. Investigate SHE.*

RF-Catcher built and coupled to BGS. \$200K.

Analyzer magnet for mass identification of SHE. \$180K.

b) *The Omnibus: unique device for heavy element and far-from-stability studies.*

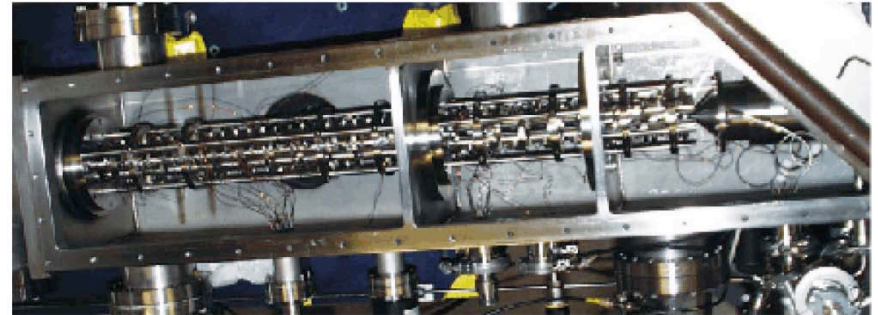
R&D for Omnibus large acceptance spectrometer.

c) *Investigations of  $\gamma$ -process reactions in inverse kinematics.*

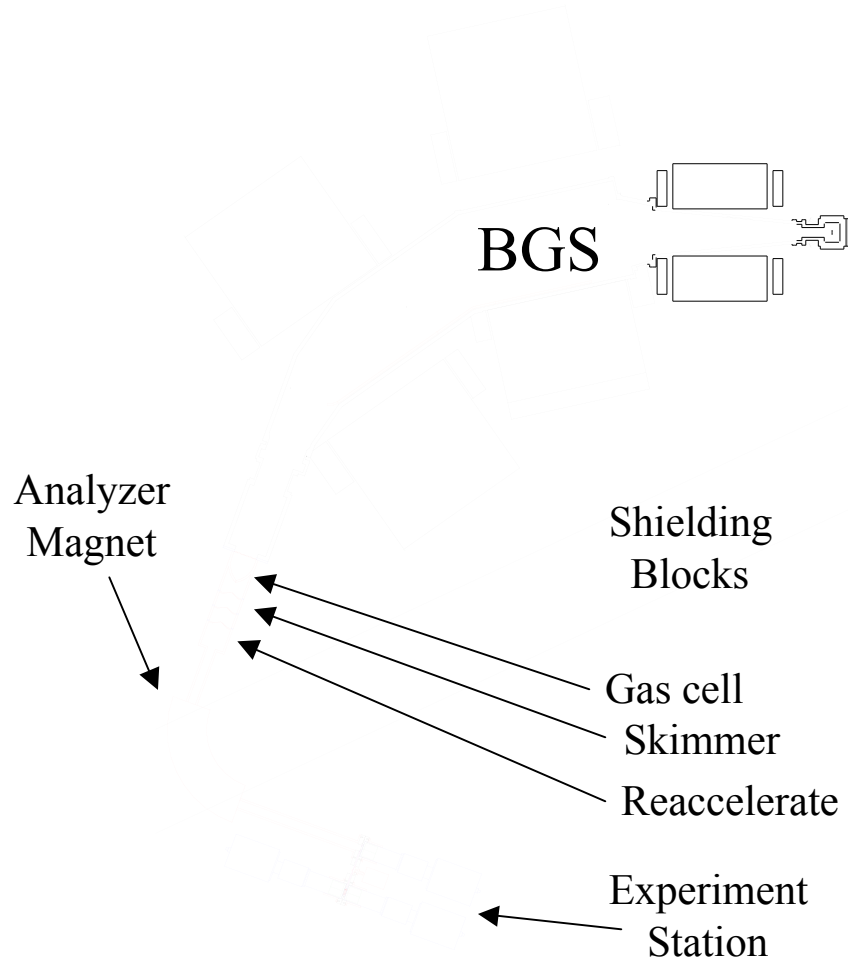
CLAIRE Stage 1 – gas-jet target using beams from the 88-Inch. \$700K.



# RF Catcher



*“The Argonne Catcher”*



- Reaction products separated from beam by the BGS
- Stopped in high-purity He gas cell – retain  $1^+$  charge
- RF field pushes ions towards exit where they are skimmed
- Re-accelerated to moderate potentials (10kV), mass separate.

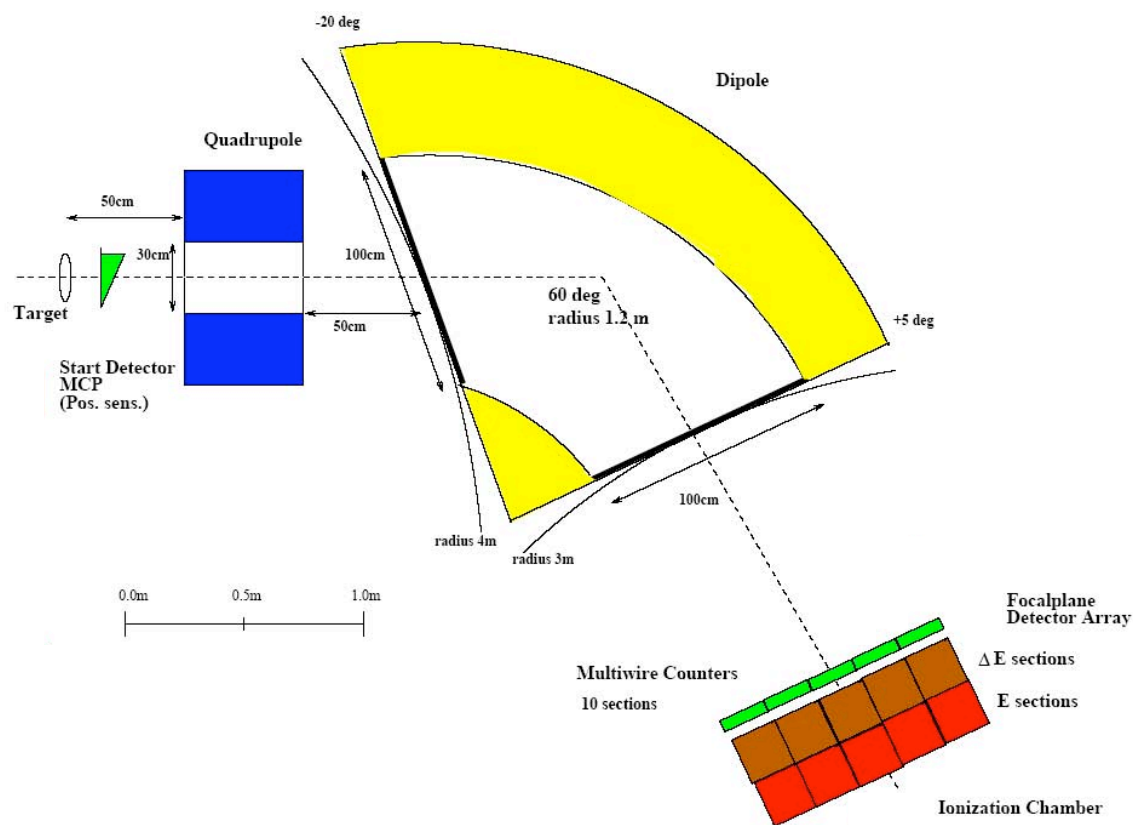
- Decay studies in low background environment
- Chemistry studies on isotopes with  $> \text{ms}$  lifetimes
- Mass identification of heaviest atoms (analyzer magnet)

# Omnibus

Concept: A superconducting, dual-mode, large-acceptance spectrometer.

Mode 1: Gas-filled “better than BGS” separator.

Mode 2: Gas-less “PRISMA-like” spectrometer.



Solid Angle  $\Omega \sim 80 \text{msr}$

Momentum Acceptance  $\sim \pm 10\%$

A Resolution  $\frac{\Delta A}{A} \sim \frac{1}{300}$

Z Resolution  $\frac{\Delta Z}{Z} \sim \frac{1}{60}$

Rotation about target  $-20^\circ \leq \theta \leq 130^\circ$

- Study nuclei far-from-stability populated in deep-inelastic reactions.
- A large-acceptance spectrometer of the type needed at future FRIB.

## GRETINA/GRETA+Omnibus: Far-From-Stability Studies

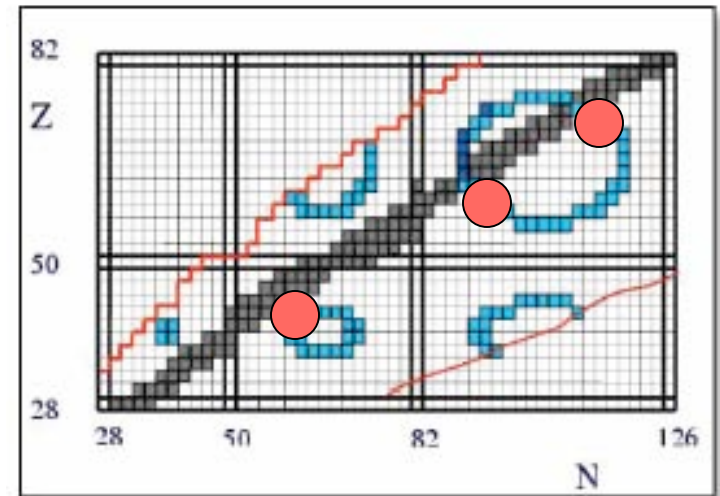
- Omnibus in gas-less mode allows Z and A identification ( $\Delta Z/Z \sim 1/80$ ,  $\Delta A/A \sim 1/300$ ) of species formed in deep-inelastic reactions (non-selective population of high-I states).
- GRETA allows us to extend reach for  $\gamma$ -spectroscopy by several more neutrons.
- Complementary to radioactive-beam studies of nuclei far-from-stability

### Some Physics Examples

- Structure of nuclei near doubly magic (Approaching  $^{78}\text{Ni}$  and  $^{132}\text{Sn}$ ).
- Stability of shell gaps in n-rich nuclei (Recent controversy in  $^{52-54}\text{Ca}$ ).
- Highly deformed states in n-rich S-Ar (Overlap of theories).
- New exotic excitation modes (Pygmy resonances in Ca isotopes)
- Shape evolution and critical-point nuclei in n-rich regions. (N-rich Zr-Mo, Ce-Nd, and Er-Yb)

$$\text{Loci of } P = \frac{N_p N_n}{N_p + N_n} \sim 5$$

Near stability this reflects rapid change from vibrational to rotational behavior

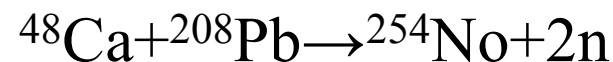


Will this hold true in n-rich nuclei?

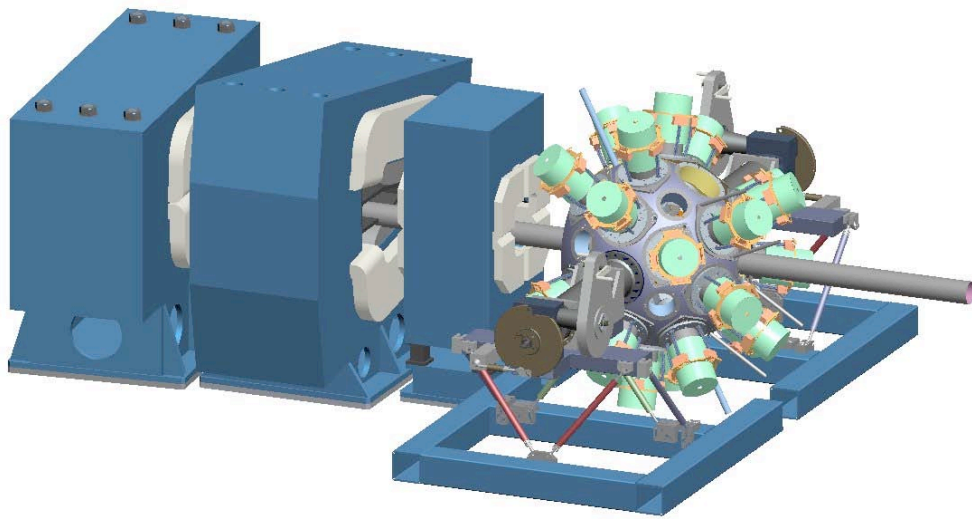
## FY11/12

- a) *High-precision mass measurements and atomic physics of heavy elements.*  
Penning trap combined with RF catcher at end of BGS. \$700K.
- b) *Transferrmium spectroscopy with most powerful set-up conceivable.*  
Gretina coupled to BGS at target position. \$250K.
- c) Construction of Omnibus. \$5000K.
- d) *High intensity ( $\sim 0.1$  A) light ion beams for H and He burning experiments.*  
CLAIRE Stage 2 – HV platform with gas-jet target. \$700K.

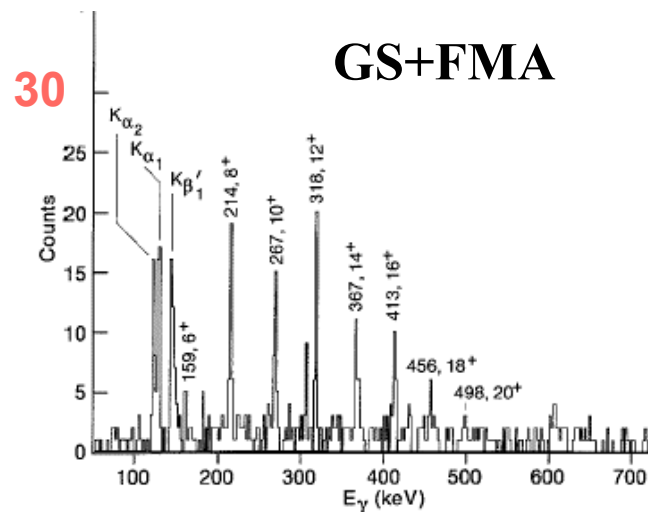
# BGS+GRETINA: Transfermium Spectroscopy



$$\sigma \sim 2 \mu\text{b}$$

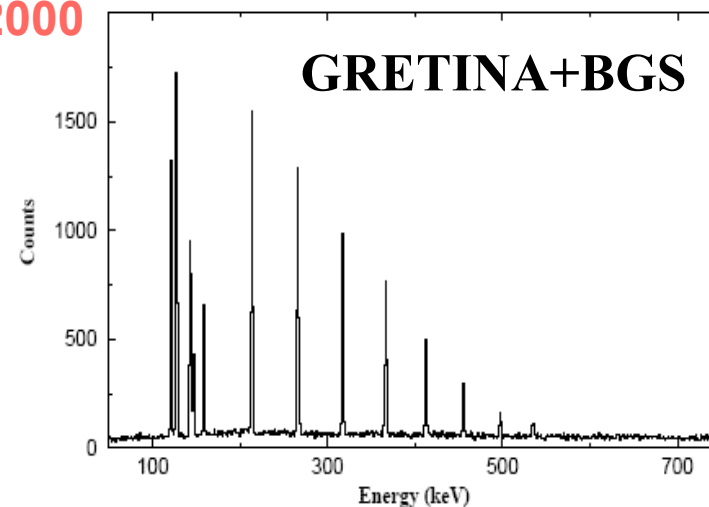


**GS+FMA**



- New generation of prompt transfermium gamma-ray spectroscopy.
- Detailed  $\gamma$ - $\gamma$  spectroscopy on even-even  $Z > 100$  nuclei possible.
- $\gamma$ -spectroscopy on odd-A nuclei
- $\gamma$ -spectroscopy of Rf ( $Z=104$ ) possible ( $^{50}\text{Ti} + ^{208}\text{Pb} \rightarrow ^{256}\text{Rf} + 2\text{n}$  with  $\sigma \sim 20$  nb).

2000



Single-particle orbits same as those in predicted spherical SHEs

## FY13/14

a) *Omnibus in gas-filled mode used for heavy element physics and chemistry.*

First scientific campaigns with Omnibus.

b) *GRETA+Omnibus in gas-less mode for DIC experiments on n-rich nuclei.*

GRETA on construction path. Host  $2\pi$  array (combined with Omnibus).

c) *R&D for CLAIRE stage 3 to study of CNO and NeNa cycles with heavier beams.*

CLAIRE Stage 3 – ECR source and RF cavities added.

## FY15/16

a) *CLAIRE is National Accelerator Facility for nuclear astrophysics*  
Add RF cavities to CLAIRE to create high intensity LINAC.

b) *Omnibus and BGS used for program of heavy element and far-from-stability studies.*

Omnibus and BGS are versatile large acceptance spectrometers.

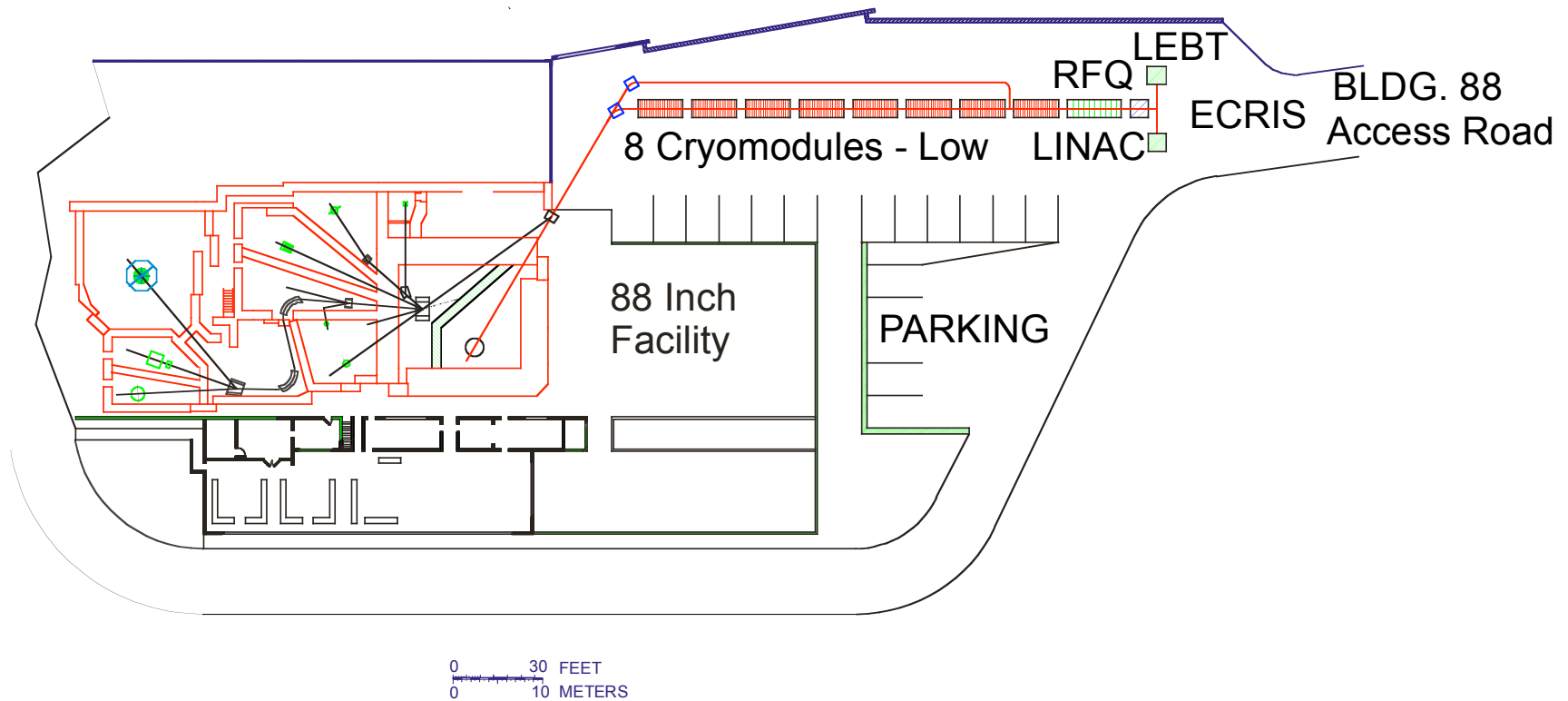
c) *Full program on hyperdeformation, transfermium spectroscopy, new excitation modes in n-rich nuclei, etc*

Host completed  $4\pi$  GRETA for stable beam campaigns.

d) *Host scientific and applied programs at 88-Inch.*

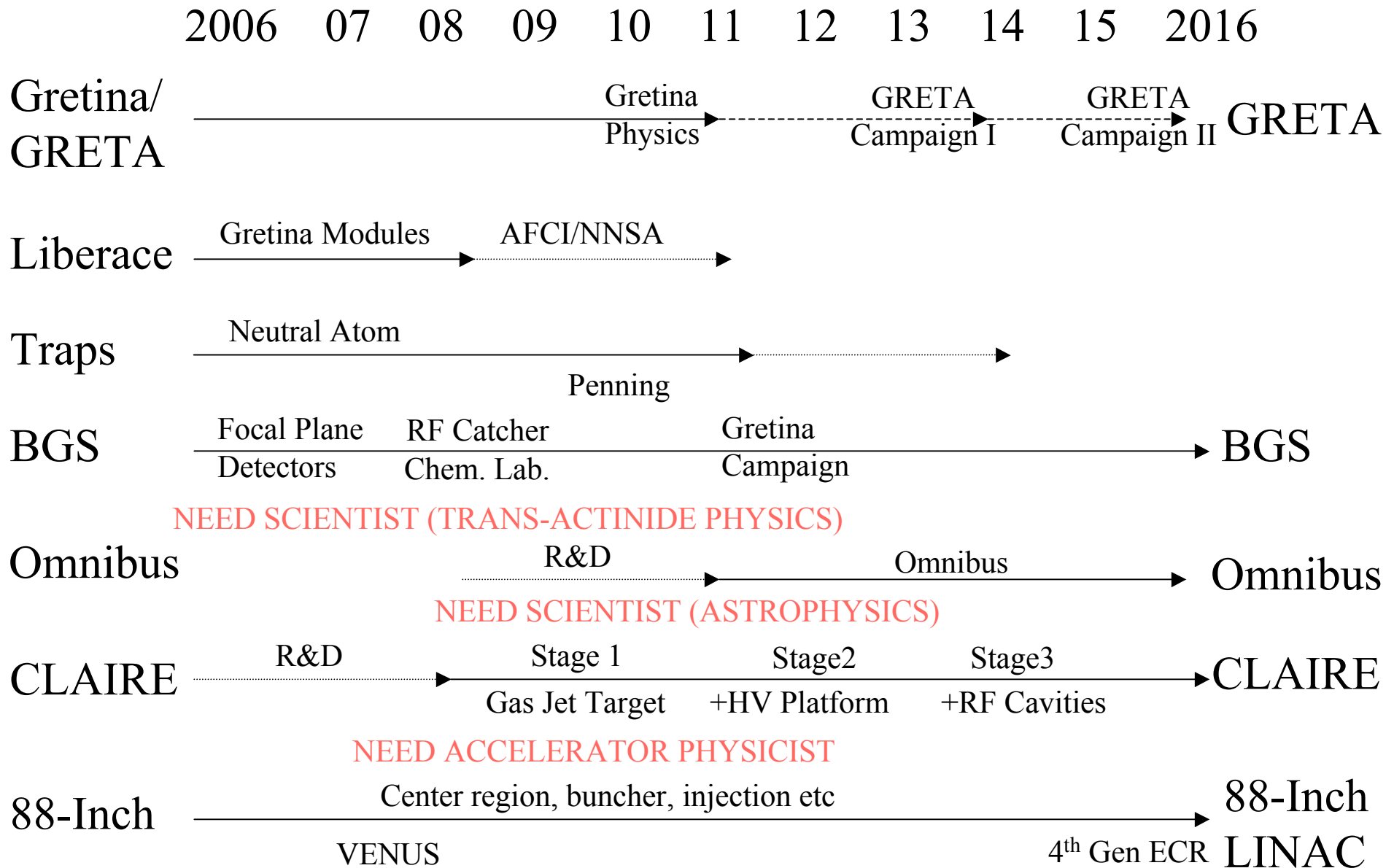
88-Inch remains versatile stable beams machine.

# A Possible Road to a High-Intensity LINAC



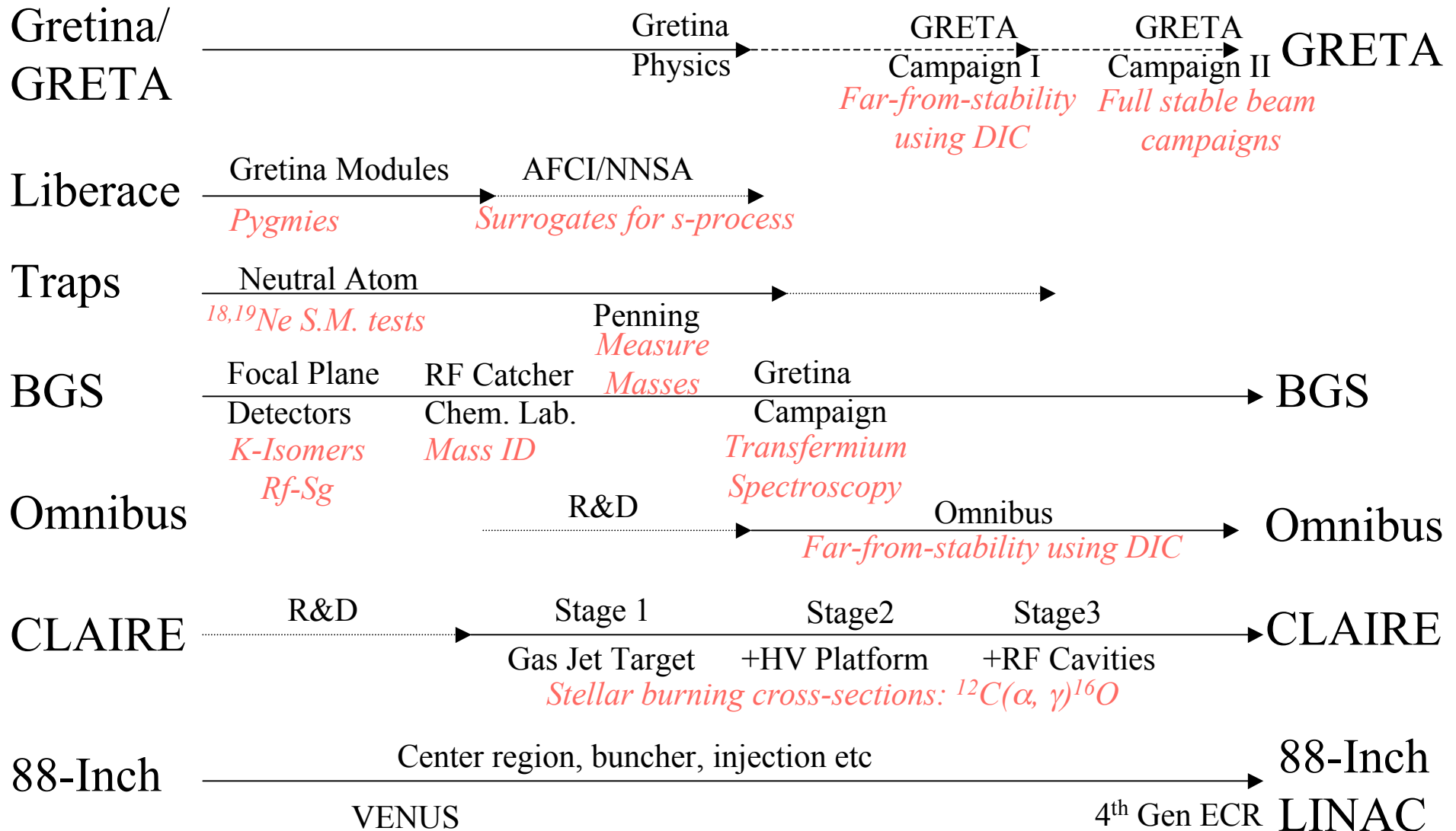


## A Plan for the Next Decade: Timeline – Growth of Manpower



## A Plan for the Next Decade: Timeline – Science Topics

2006 07 08 09 10 11 12 13 14 15 2016



## Mapping to DOE/NSAC Milestones

Year	Milestone / <b>88-Program</b> / <b>Experiment</b>
2006	<p>Measure changes in shell structure and collective modes as a function of neutron and proton number from the proton drip line to moderately neutron-rich nuclei.</p> <p><a href="#">LIBERACE program</a> using light-ion fusion-evaporation, transfer, and incomplete fusion.</p> <p>Example – lifetime measurement of excited states in <math>^{16}\text{C}</math>, the egg nucleus.</p>
2007	<p>Measure properties of the heaviest elements above <math>Z \sim 100</math> to constrain and improve theoretical predictions for super-heavy elements.</p> <p><a href="#">BGS program</a> on heavy element production, structure, and chemistry.</p> <p>Example – K-isomer studies at BGS focal plane up to Sg (<math>Z=106</math>) nuclei.</p>
2009	<p>Extend spectroscopic information to regions of crucial doubly magic nuclei far from stability such as Ni-78.</p> <p>Deep-inelastic reactions using <a href="#">GRETINA/OMNIBUS</a> and intense beams from <a href="#">VENUS</a>.</p> <p>Example – <math>^{238}\text{U} + ^{48}\text{Ca}</math> to investigate <math>N=32,34</math> sub-shells and push towards doubly magic <math>^{60}\text{Ca}</math>.</p>

## Mapping to DOE/NSAC Milestones

Year	Milestone / <b>88-Program</b> / <b>Experiment</b>
2010	<p>Complete initial measurements with the high resolving power tracking array, GRETINA, for sensitive studies of structural evolution and collective modes in nuclei.</p> <p>Use <u>GRETINA</u> modules and full <math>1\pi</math>-array for physics campaigns at <u>88-Inch</u>.</p> <p>Example – pygmy resonance structure in <math>^{208}\text{Pb}</math> from H.I. inelastic scattering.</p> <p>Example – transfermium spectroscopy (prompt) up to Rf (Z=104) using BGS.</p>
2010	<p>Reduce uncertainties of the most crucial stellar evolution nuclear reactions (e.g. <math>^{12}\text{C}(\alpha, \gamma)^{16}\text{O}</math>) by a factor of two, and others (e.g., MgAl cycle) to limits imposed by accelerators and detectors.</p> <p>Development and use of <u>CLAIRE</u> (Stages 2 and 3)</p> <p>Example – direct measurement of <math>^{12}\text{C}(\alpha, \gamma)^{16}\text{O}</math>.</p>
2011	<p>Measure neutron capture reactions, including s-process branch-point nuclei, to constrain s-process isotopic abundances.</p> <p>Surrogate techniques developed with <u>LIBERACE</u> for AFC studies (currently, deducing n-induced actinide cross-sections reactions).</p> <p>Example – <math>^{153}\text{Gd}(n, \gamma)</math> determined from surrogates such as <math>^{154}\text{Gd}(\alpha, \alpha')</math> and <math>^{155}\text{Gd}(^3\text{He}, \alpha)</math></p>

## Plan Rests Heavily on 88-Inch Operations

- Current funding for 88-Inch will mean rapid end to science program at the accelerator.
- Requested funding brings us to “sustainable operation”.
- Scientific plan requires upgrades and improvements.  
(e.g. coupling VENUS to 88-Inch to give full range of intense heavy-ion beams for heavy element research).
- Projects like CLAIRE and Omnibus require increased manpower and technical support (accelerator physicist, engineers, techs...)
- We want to host “National Facility”-scale scientific programs.  
(GRETINA/GRETA, BGS/Omnibus, CLAIRE...)

## Summary

- Developed a strong scientific vision for the next decade
- Aligned with current priorities of the National Low Energy Program
- 88-Inch is central to the plan
- Providing an essential stable beam capability in the U.S.
  - Vital to the BASE and applied programs
  - Unique heavy element research effort
  - New opportunities in nuclear structure
  - Training the next generation
- Developing new facilities and apparatus for U.S. nuclear science
  - GRETINA/GRETA
  - BGS/Omnibus
  - CLAIRE



“Give us the tools and we will finish the job!”